Symmetry reductions in loop quantum gravity

based on classical gauge fixings

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Plan of the talk



- 2 General strategy
- 3 Bianchi I: Details on classical derivation
- 4 Bianchi I: Details on quantum theory
- 5 Spherical symmetry (sketch)



Outline



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- Mini / midi-superspace quantisation
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 - Schwarzschild black hole [Kastrup, Thiemann '93; Kuchař '94, Gambini, Pullin '13]
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 - ▶ Quantisation ↔ reduction [Bojowald '04; Engle '05; Hanusch '13]
 - Embedding of states [Engle '07; Brunnemann, Fleischhack '07; Fleischhack '10]

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- Code symmetry as f(p,q) = 0, impose $f(p,q) |\Psi\rangle_{sym} = 0 \leftarrow \text{this talk}$
 - Bianchi I models [NB '14]
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Recipe for canonical loop quantisation

Hamiltonian connection formulation of a gravitational theory

 $\text{Metric} \rightarrow \text{connection}$ with gauge symmetry under a local gauge group $_{(e.g. SU(2))}$

+ certain mathematical properties (avoid with new vacua? [Dittrich, Geiller '14; Bahr, Dittrich, Geiller '15])

- compact gauge group
- real variables
- canonical Poisson brackets
- only first class constraints for Dirac quantisation (otherwise e.g. Gupta-Bleuler as in EPRL model, or Master constraint)

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Methods of **Algebraic quantum field theory**

(C^* -algebras, GNS-construction, ...)

Kinematically quantised gravitational theory

[Rovelli, Smolin, Ashtekar, Isham, Lewandowski, Marolf, Mourao, Thiemann, Sahlmann, ...]

- Regularise Hamiltonian constraint [Thiemann '96; ...]
- Compute quantum constraint algebra [Thiemann '96; ...; Varadarajan, Laddha et al. '11-]

Suitable classical starting point

- Gauge fix spatial diffeomorphisms adapted to the symmetry reduction
- ▶ Go to the reduced phase space, i.e. solve constraints or employ Dirac bracket
- Find new connection variables on Γ_{red} (not Ashtekar-Barbero variables)

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- Find phase space functions $f_i(p,q) = 0$ in the symmetric subspace
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(3) Impose reduction conditions $f_i = 0$ as operator equations: $\hat{f}_i |\Psi\rangle_{sym} = 0$

- Find subspace of quantum reduced states $|\Psi\rangle_{svm}$
- Find observables $\hat{\mathcal{O}}_{sym}$ w.r.t. reduction constraints: $[\hat{\mathcal{O}}_{sym}, \hat{f}_i] = 0$

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§ Relate observables $\hat{\mathcal{O}}_{sym}$ to mini- / midisuperspace parameters

- Map observables
- Study dynamics

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Gauge fixing to obtain suitable coordinates

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- **(3)** Coordinatise the reduced phase space via $q_{xx}, q_{yy}, q_{zz}, P^{xx}, P^{yy}, P^{zz}$
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Choose connection type variables

- 1 Define $e_a e_a = q_{aa}$, $e_a e^a = 1$, without summation, and $E^a = \sqrt{\det q} e^a$
- 2 Define $K_a = K_{ab}e^b$ with K_{ab} being the extrinsic curvature constructed form P^{ab}
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- *K_a*, *E^b* are like Ashtekar-Barbero variables without internal indices ⇒ Abelian gauge theory (Poisson bracket of Maxwell theory)

At this stage, only Hamiltonian constraint and reduced spatial diffeomorphisms left.

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Result:

Direct consequences of a Bianchi I reduction can be imposed as spatial diffeomorphisms and a Gauß law on the (quantised) reduced phase space (as operator equations).

Norbert Bodendorfer (Univ. of Warsaw)

Symmetry reductions in LQG

Classical preparations III: Summary

Phase space: (full GR admitting diagonal metric gauge)

1
$$K_a(\sigma)$$
, $E^b(\sigma)$ are 3 + 3 canonical variables per spatial point σ

- 2 Remaining constraints are
 - reduced spatial diffeomorphisms (preserving the diagonal gauge)
 - 2 Hamiltonian constraint

Direct consequences of a reduction to Bianchi I are

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Strategy:

Quantise full phase space via LQG techniques

2 Impose symmetry reduction by imposing $\tilde{C}_a = 0 = G$ at the quantum level

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Standard LQG type quantisation

• Compute holonomies $h_{\gamma}^{\lambda}(K) := \exp\left(i\lambda \int_{\gamma} K_{a} ds^{a}\right)$ and fluxes $E(S) = \int_{S} E^{a} d^{2}s_{a}$ γ path, S surface, $\lambda \in \mathbb{Z}$ for U(1), or $\lambda \in \mathbb{R}$ for \mathbb{R}_{Bohr} see e.g. [Corichi, Krasnov '97] for U(1)

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- ② Define positive linear Ashtekar-Lewandowski functional on holonomy-flux algebra
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Remarks

• For \mathbb{R}_{Bohr} : $\lim_{R\to\infty} \frac{1}{2R} \int_{-R}^{R} dx f(x) = \int_{\mathbb{R}_{Bohr}} d\mu_{H} f(x)$ provides **normalised** and translation invariant Haar measure \Rightarrow per edge: $\mathcal{H} = L^{2}(\mathbb{R}_{Bohr}, d\mu_{H})$

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• Choosing $\lambda \in \mathbb{Z}$ over $\lambda \in \mathbb{R}$ (i.e. compactifying $\int_{\gamma} K_a ds^a$) has no justification at this stage (also not later)

Quantisation II: Area operator

Area operator for Abelian theory

- $A(S) = |E(S)| = |\int_S E^a d^2 s_a|$ is analogous to (absolute value of) electric flux
- Important difference to non-Abelian, e.g. SU(2), area op. $\int_{S} \sqrt{|E^{i}E_{i}|}$:
 - Absolute value is outside of the integral
 - E(S) does not detect closed contractible loops for closed S

While one can also define "non-Abelian like" area operator here, the Abelian one will turn out to be most useful.



Quantisation II: Area operator

Area operator for Abelian theory

- $A(S) = |E(S)| = |\int_S E^a d^2 s_a|$ is analogous to (absolute value of) electric flux
- Important difference to non-Abelian, e.g. SU(2), area op. $\int_{S} \sqrt{|E^{i}E_{i}|}$:
 - Absolute value is outside of the integral
 - \triangleright E(S) does not detect closed contractible loops for closed S

While one can also define "non-Abelian like" area operator here, the Abelian one will turn out to be most useful.

- Non-trivial topology: A(S) can detect Wilson loops even for closed S



 $\hat{A}(S)$ measures intersection number $N_{\text{int}} imes$ rep. label: $\hat{A}(S) \left| h_{\gamma}^{\lambda} \right\rangle = \left| N_{\text{int}} \lambda \right| \left| h_{\gamma}^{\lambda} \right\rangle$

Quantisation III: Imposing the symmetry reduction

Reduction constraints are very familiar from full theory

- **1** All spatial diffeomorphisms: $\tilde{C}_a[N^a] = \int_{\Sigma} d^3 \sigma E^a \mathcal{L}_{\vec{N}} K_a = 0$
- 2 Abelian Gauß law: $G[\omega] = \int_{\Sigma} d^3 \sigma \, \omega \, \partial_a E^a = 0$

⇒ spatially diffeomorphism invariant and gauge invariant charge (spin) networks!

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Simplest choice of quantum state Consider spin network made from 3 Wilson loops wrapping around $\mathbb{T}^1_x, \mathbb{T}^1_y, \mathbb{T}^1_z$, meeting in a single vertex *v*. [c.f. Husain '91, '05] Mapping to Bianchi I LQC states of [Ashtekar, Wilson-Ewing '09] $|\lambda_x, \lambda_y, \lambda_z\rangle \mapsto |p_1, p_2, p_3\rangle$



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Observables w.r.t. the reduction constraints

Area of closed surfaces → 3 non-trivial areas A(T²_x), A(T²_y), A(T²_z)
 Diff-equiv. classes of Wilson loops → 3 non-trivial closed loops along T¹_x, T¹_y, T¹_z

Hamiltonian constraint / true Hamiltonian (via deparametrisation)

Take original Hamiltonian:

- Evaluate at $q_{a\neq b} = 0$ because of gauge fixing
- Discard $P^{a\neq b}$, $\partial_a e_b$, and $\partial_a K_b$ terms because of reduction constraints

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Regularise constraint operator (graph preserving)

- Substitute e_a either by fluxes or Thiemann's trick $e_a = 2\{K_a, V\}$
- Approximate K_a via holonomies: $\int K_a ds^a \approx \sin(\lambda \int K_a ds^a)/\lambda$

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Choice of λ is crucial! U(1) vs. \mathbb{R}_{Bohr}

• U(1) choice: $\lambda = 1$ gives best approximation \Rightarrow "old" LQC dynamics [Ashtekar, Bojowald, Lewandowski '03]

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• \mathbb{R}_{Bohr} allows arbitrarily small $\lambda \in \mathbb{R}$ for better approximation. "improved" LQC choice: $1/\lambda_x = \sqrt{|E^y E^z / E^x|} = \text{size of universe in } x$ -direction \Rightarrow "new" LQC dynamics [Ashtekar, Pawlowski, Singh '06; Ashtekar, Wilson-Ewing '09]

Outline

Approaches to symmetry reductions

- 2 General strategy
- 3 Bianchi I: Details on classical derivation
- 4 Bianchi I: Details on quantum theory
- 5 Spherical symmetry (sketch)

Conclusion

ADM phase space in radial gauge (without details here)

Start with ADM phase space

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- Start with ADM phase space
- 2 Define adapted coordinate system

[Duch, Kamiński, Lewandowski, Świeżewski '14]



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3 Impose radial gauge $q_{ra} = \delta_{ra}$



ADM phase space in radial gauge (without details here)



 $\begin{array}{l} \overline{A}, \overline{B} = \theta, \phi, \quad \{A_A^i(\sigma), E_j^B(\sigma')\} = \delta^{(3)}(\sigma, \sigma')\delta_A^B \delta_j^i \\ A_A^i, \overline{E}_i^B = \text{variables of 3d gravity, with spatial slice } S_r^2 \end{array}$



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Find conditions compatible with spherical symmetry

P^{rA} = 0 ⇔ generator of all spatial diffeomorphisms preserving all S²_r (Follows form non-existence of non-zero spherically symmetric vector field on S²)

Impose spherical symmetry as invariance under S_r^2 -preserving diffeomorphisms.

(These are active diffeomorphisms with respect to the (r, θ, ϕ) coordinate system)

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Symmetry reductions in LQG

Perform standard LQG-type quantisation (roughly similar to lattice field theory)

() SU(2) gauge theory with holonomies restricted to lie in an S_r^2

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Observables w.r.t. the reduction constraints

Areas of the S²_r → 4πR(r)² := ∫_{S²_r} d²x √det q_{AB}
 Averaged trace of momenta → P_R(r) := ²/_{R(r)} ∫_{S²_r} d²x P^{AB} q_{AB}

(+ all other S_r^2 -preserving diffeomorphism invariant observables. \Rightarrow more than in classically reduced theory!)

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What about dynamics?

More challenging than for Bianchi I, ongoing work with A. Zipfel. First steps

- $\bullet~$ Map states in classically reduced \rightarrow quantum reduced theory
- Compute quantum algebra $[\hat{R}(r), \hat{P}_{R}(r')]$ from full theory

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Suitable classical starting point

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- Sph. sym.: ADM in radial gauge
- Identification of constraints imposed by symmetry reduction
 - Bianchi I: all spatial diffeomorphisms and Abelian Gauß law
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- **Quantise** the reduced phase space via LQG techniques
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- **§ Relate observables** w.r.t. reduction constraints to mini- / midisuperspace
 - Bianchi I: three areas and conjugate momenta, dynamics agree with mini-superspace quantisation
 - Sph. sym.: R(r) and $P_R(r)$, dynamics under investigation

6 Future work: perturbations to Bianchi I, coarse graining, spherical collapse...

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Thank you for your attention!