

# Purity is not eternal at the Planck scale

Michele Arzano

Dipartimento di Fisica  
"Sapienza" University of Rome



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# 40 Years of Hawking radiation

*Nature Vol. 248 March 1 1974*

## Black hole explosions?

QUANTUM gravitational effects are usually ignored in calculations of the formation and evolution of black holes. The justification for this is that the radius of curvature of space-time outside the event horizon is very large compared to the Planck length  $(GH/c^3)^{1/2} \approx 10^{-33}$  cm, the length scale on which quantum fluctuations of the metric are expected to be of order unity. This means that the energy density of particles created by the gravitational field is small compared to the space-time curvature. Even though quantum effects may be small locally, they may still, however, add up to produce a significant effect over the lifetime of the Universe  $\approx 10^{17}$  s which is very long compared to the Planck time  $\approx 10^{-43}$  s.

...

S. W. HAWKING

*Department of Applied Mathematics and Theoretical Physics  
and  
Institute of Astronomy  
University of Cambridge*

Received January 17, 1974.

- <sup>1</sup> Bardeen, J. M., Carter, B., and Hawking, S. W., *Commun. math. Phys.*, **31**, 161-170 (1973).
- <sup>2</sup> Hawking, S. W., *Mon. Not. R. astr. Soc.*, **152**, 75-78 (1971).
- <sup>3</sup> Penrose, R., in *Relativity, Groups and Topology* (edit. by de Witt, C. M., and de Witt, B. S). Les Houches Summer School, 1963 (Gordon and Breach, New York, 1964).
- <sup>4</sup> Hawking, S. W., and Ellis, G. F. R., *The Large-Scale Structure of Space-Time* (Cambridge University Press, London 1973).
- <sup>5</sup> Hawking, S. W., in *Black Holes* (edit. by de Witt, C. M., and de Witt, B. S), Les Houches Summer School, 1972 (Gordon and Breach, New York, 1973).
- <sup>6</sup> Beckenstein, J. D., *Phys. Rev.*, **D7**, 2333-2346 (1973).

## The information paradox in a nutshell

**The stage:** free quantum field on a Schwarzschild black hole background

Essence of Hawking effect: vacuum state for a **free falling observer**  $|0\rangle \neq$   
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- Black hole completely evaporates  $\equiv$  no horizon, no “inside” region
- The mixed state  $\rho$  cannot be a partial trace of a pure state since there's **no degrees of freedom to trace out** left!

Is purity eternal?

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15 NOVEMBER 1976

### **Breakdown of predictability in gravitational collapse\***

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Commun. Math. Phys. 87, 395–415 (1982)

### The Unpredictability of Quantum Gravity

S. W. Hawking

University of Cambridge, D.A.M.T.P., Cambridge CB3 9EW, England

*“I present a number of axioms that the asymptotic Green functions should obey in any reasonable theory of quantum gravity. These axioms are the same as for ordinary quantum field theory in flat spacetime, except that one axiom, that of asymptotic completeness, is omitted. This allows pure quantum states to decay into mixed states.”*



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are preserved by time evolution they (re)-discovered the **Lindblad equation**

$$\dot{\rho} = -i[H, \rho] - \frac{1}{2} h_{\alpha\beta} \left( Q^\alpha Q^\beta \rho + \rho Q^\beta Q^\alpha - 2Q^\alpha \rho Q^\beta \right)$$

$h_{\alpha\beta}$  is a hermitian matrix of constants and  $Q^\alpha$  form a basis of hermitian matrices

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### ABSTRACT

Motivated by Hawking's proposal that the quantum-mechanical density matrix  $\rho$  obeys an equation more general than the Schrödinger equation, we study the general properties of evolution equations for  $\rho$ . We argue that any more general equation for  $\rho$  violates either locality or energy-momentum conservation.

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Lindblad time evolution is still problematic since: “[...] *loss of purity is incompatible with the weakest possible form of Lorentz covariance.*”

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**MA: 1403.6457; Phys. Rev. D 90, 024016 (7 July 2014)**

## Outline

- **Topological particles and curved momentum space in 3d gravity**
- **Quantum double, deformed symmetries and Lindblad evolution**
- **4d case: de Sitter momentum space and  $\kappa$ -deformed symmetries**
- **Deformed Lindblad evolution from  $\kappa$ -Poincaré**
- **Conclusions and outlook**

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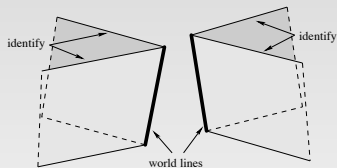
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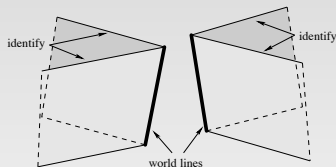
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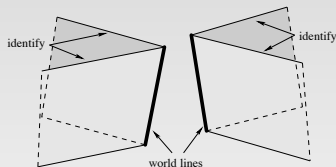
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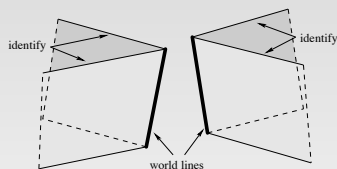
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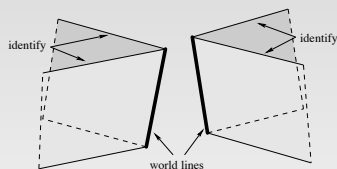
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**Momenta of particles coupled to 3d gravity = elements of a non-abelian group!**

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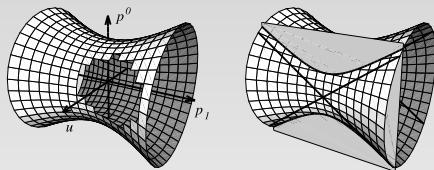
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The unit determinant condition  $u^2 + p^2/\kappa^2 = 1 \implies$

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The unit determinant condition  $u^2 + p^2/\kappa^2 = 1 \implies$



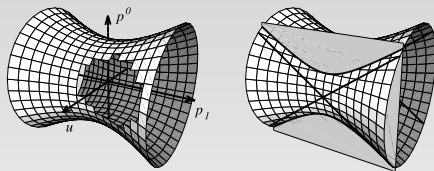
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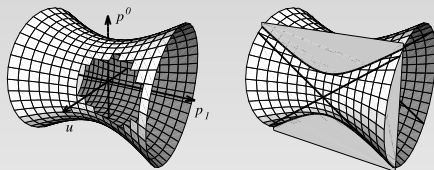


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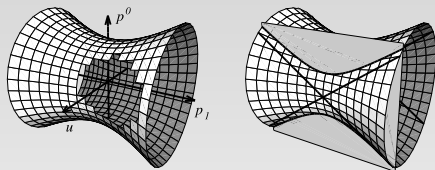
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- **Lorentz transformation:**  $h' = ghg^{-1}$ , undeformed on  $p^\mu$  e.g. boost in the 1-direction  $g = e^{\frac{1}{2}\eta\gamma_2}$

$$\begin{cases} p'^0 = p^0 \cosh \eta - p^1 \sinh \eta \\ p'^1 = p^1 \cosh \eta - p^0 \sinh \eta \\ p'^2 = p^2 \end{cases} .$$

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- ii) momentum coordinates obey a non abelian composition rule indeed

$$p_\mu \oplus q_\mu = v(q) p_\mu + u(p) q_\mu + \frac{1}{\kappa} \epsilon_{\mu\nu\sigma} p^\nu q^\sigma = p_\mu + q_\mu + \frac{1}{\kappa} \epsilon_{\mu\nu\sigma} p^\nu q^\sigma + \mathcal{O}(1/\kappa^2) \neq q_\mu \oplus p_\mu$$

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what does this mean physically?

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## Basic quantum theory

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consider e.g. a projector  $\pi_k = |k\rangle\langle k|$  seen as the “outer product” of a ket and a bra namely  $\pi_k = \pi(|k\rangle \otimes \langle k|) = |k\rangle\langle k|$

$$P_\mu(\pi_k) = \pi(P_\mu(|k\rangle \otimes \langle k|)) = \pi(P_\mu |k\rangle \otimes \langle k| - |k\rangle \otimes \langle k| P_\mu) = [P_\mu, \pi_k]$$

i.e. just the familiar **adjoint action**

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**Key point:** the action on operators will be deformed accordingly

## Deformed translations and Lindblad evolution in three dimensions

The deformed translation generators of  $\mathcal{D}(SL(2, \mathbb{R}))$  in the “cartesian” basis  $P_\mu$ :

$$\Delta P_\mu = P_\mu \otimes \mathbb{1} + \mathbb{1} \otimes P_\mu + \frac{1}{\kappa} \epsilon_{\mu\nu\sigma} P^\nu \otimes P^\sigma, \quad S(P_\mu) = -P_\mu.$$

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which can be rewritten in Lindblad form as

$$\dot{\rho} = -i[P_0, \rho] - \frac{1}{2} h_{ij} \left( P^i P^j \rho + \rho P^j P^i - 2P^j \rho P^i \right)$$

with “dissipation” matrix is given by

$$h = \frac{i}{\kappa} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & -1 & 0 \end{pmatrix}$$



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**In our case  $h$  is not positive definite nor real**

Further work needed to establish properties of our Lindblad evolution...

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The non-abelian composition of momenta in "flat slicing" or **bicrossproduct** coordinates

$$p_0(k_0, \mathbf{k}) = \kappa \sinh k_0/\kappa + \frac{\mathbf{k}^2}{2\kappa} e^{k_0/\kappa},$$

$$p_i(k_0, \mathbf{k}) = k_i e^{k_0/\kappa},$$

$$p_4(k_0, \mathbf{k}) = \kappa \cosh k_0/\kappa - \frac{\mathbf{k}^2}{2\kappa} e^{k_0/\kappa}.$$

reads  $k \oplus l = (k^0 + l^0; k^j e^{-\frac{l^0}{\kappa}} + l^j)$

The non-abelian composition of momenta reflects a **non-Leibniz action** of *spatial* translation generators

$$K_i(\mathbf{e}_{\mathbf{k}_1} \otimes \mathbf{e}_{\mathbf{k}_2}) = K_i(\mathbf{e}_{\mathbf{k}_1}) \otimes \mathbf{e}_{\mathbf{k}_2} + \exp(-K_0/\kappa)(\mathbf{e}_{\mathbf{k}_1}) \otimes K_i(\mathbf{e}_{\mathbf{k}_2})$$



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in the limit  $\kappa \rightarrow \infty$  recover ordinary Poincaré algebra

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a straightforward calculation leads to a *non-symmetric* **Lindblad equation**

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- however the quantum adjoint actions of  $N_i$  and  $P_0$  satisfy

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in this sense the  $\kappa$ -Lindblad equation follows a **deformed notion of covariance**

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  - ▶ *phenomenology* of  $\kappa$ -Lindblad evolution? (Ellis et al. “Search for Violations of Quantum Mechanics,” Nucl. Phys. B **241**, 381 (1984) and following works)

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- and finally application to **black hole quantum evolution**...

